

*Acceleration of the spherically symmetric Boltzmann equation
for supernova explosion simulations
on GPU, Pezy-SC, and SX-Aurora*

Hideo Matsufuru

High Energy Accelerator Research Organization, KEK



Kohsuke Sumiyoshi

National Institute of Technology, Numazu College

Multi-dimensional Modeling and Multi-Messenger observation
from Core-Collapse Supernovae (4M-COCOS)

Fukuoka Univ., Fukuoka, Japan, 21-24 October 2019



Core-collapse supernovae



- **Supernova explosion simulation**
 - Matter: hydrodynamics
 - Neutrino transport is important: described by Boltzmann equation
 - Equation of state for high temperature and density
 - Gravitational effect governed by general relativity
- **Large scale simulation is essential to understand explosion mechanics**
 - Fluid instability may play a key role (dimensionality is important)
 - **Boltzmann eq.: evolution eq. of 6D distribution function**
$$\frac{dx^\mu}{d\tau} \frac{\partial f_\nu}{\partial x^\mu} + \frac{dp^\mu}{d\tau} \frac{\partial f_\nu}{\partial p^\mu} = \left(\frac{\delta f_\nu}{\delta \tau} \right)_{\text{coll}} \quad f_\nu(x, p) : \text{neutrino distribution function}$$
 - Quantitative (spatial) 3D simulation is only possible at Exa-scale
 - Systematic study is needed for 1D, 2D, 3D simulations
- **Numerical simulations**
 - Massively parallel supercomputers (K, Fugaku, ...)
 - **How about accelerators ?**



Spherically symmetric system

- Spherically symmetric system

- Systematic survey of massive stars is necessary
- Basis for 2D/3D simulations and observations
- 1st principle calculation (Full GR+Hydro.+Boltzmann)
- Testbed for developing GPU code

- Numerical setup

- GR Lagrangian hydrodynamics + S_N + Implicit scheme
- Boltzmann equation for neutrinos
$$\frac{\partial f_\nu}{\partial t} + \mu \frac{\partial f_\nu}{\partial r} + \frac{(1 - \mu^2)}{r} \frac{\partial f_\nu}{\partial \mu} = \left(\frac{\delta f_\nu}{\delta t} \right)_{\text{coll}}$$
- Linear equation is solved at every step of time evolution
 - Primary bottleneck
 - BiCGStab algorithms for a block tridiagonal matrix
- Secondary bottlenecks:
 - Collision term in Boltzmann equation
 - Inversion of diagonal block matrices

- These bottlenecks are target of offloading to GPU/Pezy-SC



Machines

Performance is measured on the following machines

- NVIDIA P100 GPU (Pascal)

H.M. & K.S. Lecture Note in Comp. Sci. 10962 (2018) 440

H.M. & K.S. IEEE Xplore, Proceedings of CANDAR2018, 56

- 4.7 TFlops/device for double precision
- High memory and host-device bandwidths

- Pezy-SC: Siren-Blue @KEK

- Many-core accelerator
- 1.5 TFlops/device
- 1024 MIMD cores with hierarchical structure
- Liquid cooling for low power consumption

- NEC SX-Aurora TSUBASA

- Vector architecture: Latest product of SX-series
- 2.46 TFlops/Vector Engine
- High memory bandwidth: 1228 GB/s (B/F=0.5)
- Vector Engine works as a usual MPI node

